

P423: Lecture 8: Zeeman and hyperfine effects

Zeeman effect

Effect of B field on H-atom. To split the states, apply a B-field!

$$H_Z = - (\vec{\mu}_{orb} + \vec{\mu}_{spin}) \cdot \vec{B}_{ext}$$

the
$$\begin{cases} \vec{\mu}_{spin} = -\frac{e_0}{m} \vec{S} \\ \vec{\mu}_{orb} = -\frac{e_0}{2m} \vec{L} \end{cases} \quad (g_0=2)$$

$$H_Z = \frac{e_0}{2m} (\vec{L} + 2\vec{S}) \cdot \vec{B}_{ext}$$

To find energy levels, of H-atom, need to consider "internal field" due to spin orbit coupling:

$$H_{so} = -\vec{\mu}_{spin} \cdot \vec{B}_{int}$$
$$\vec{B}_{int} = \frac{1}{4\pi\epsilon_0} \frac{e_0}{mc^2 a^3} \vec{L}$$

Two approximation regimes depending on strength of B_{ext}

$|\vec{B}_{ext}| \ll |\vec{B}_{int}| \Rightarrow H_Z$ is a perturbation on l -states

$|\vec{B}_{ext}| \gg |\vec{B}_{int}| \Rightarrow (H_n^1 + H_n^2)$ is a perturbation on $|l, m\rangle$ or $|\pm \frac{1}{2}\rangle$ states.

2)

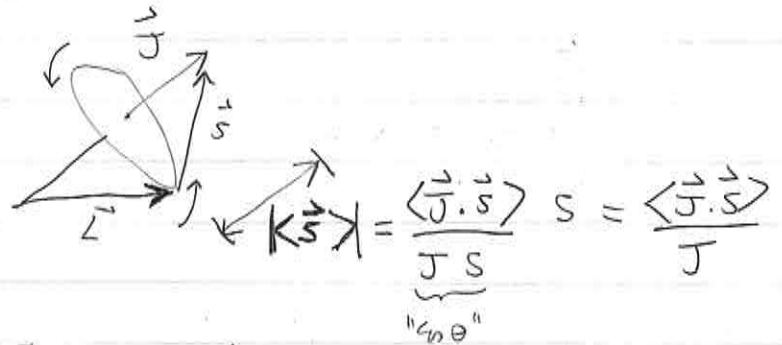
For Hydrogen, $B_{ext} \approx 12 T$ (very strong!) (assuming $|\vec{L}| \sim \hbar$, $\mu \sim \mu_B$)

a) weak field case:

$$E_B^1 = \langle j m_j | H'_B | j m_j \rangle = \frac{e_0}{2m} B_{ext} \cdot \langle \vec{L} + 2\vec{S} \rangle$$

How to calculate $\langle \vec{L} + 2\vec{S} \rangle = \underbrace{\langle \vec{J} \rangle}_{\text{easy!}} + \underbrace{\langle \vec{S} \rangle}_{?}$ in the j basis?

Use Classical intuition:



Also, $\langle \vec{S} \rangle \parallel \langle \vec{J} \rangle$

$$\Rightarrow \langle \vec{S} \rangle = \frac{\langle \vec{J} \cdot \vec{S} \rangle}{J} \frac{\langle \vec{J} \rangle}{J} = \frac{\langle \vec{J} \cdot \vec{S} \rangle}{\langle J^2 \rangle} \langle \vec{J} \rangle$$

But $\vec{L} = \vec{J} - \vec{S} \Rightarrow L^2 = J^2 - 2\vec{J} \cdot \vec{S} + S^2$

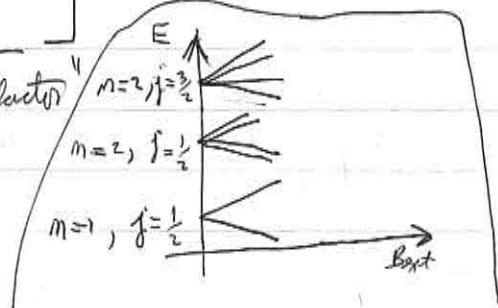
$$\Rightarrow \vec{J} \cdot \vec{S} = \frac{1}{2}(J^2 + S^2 - L^2) \Rightarrow \langle \vec{J} \cdot \vec{S} \rangle = \frac{\hbar^2}{2} [j(j+1) + s(s+1) - l(l+1)]$$

$$\Rightarrow \langle \vec{S} \rangle = \frac{\frac{\hbar^2}{2} [j(j+1) + \frac{3}{4} - l(l+1)]}{j(j+1)\hbar^2} \langle \vec{J} \rangle$$

$$\Rightarrow \langle \vec{L} + 2\vec{S} \rangle = \left[1 + \frac{j(j+1) - l(l+1) + \frac{3}{4}}{2j(j+1)} \right] \langle \vec{J} \rangle$$

g_J "Landé g-factor"

$$\Rightarrow E_B^1 = g_J \mu_B B_{ext} m_j$$



(b) Strong field case:

$$H_D = \frac{e_0}{2m} (\vec{L} + 2\vec{S}) \cdot \vec{B}_{ext} \quad \text{with } |H_D| \gg |H_{relativistic}|$$

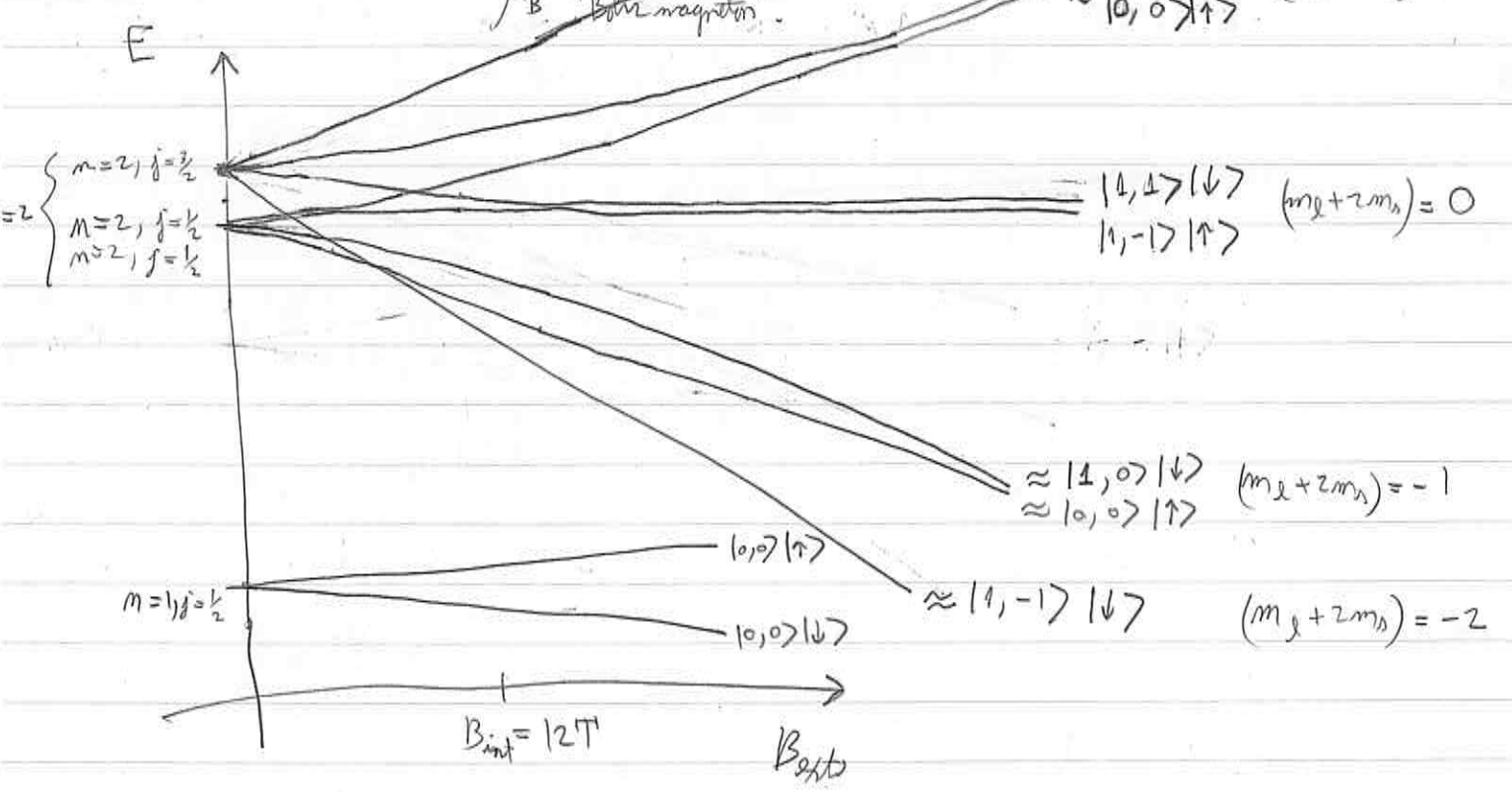
⇒ Use $|l, m_l\rangle |1/2, m_s\rangle$ basis!

For $\vec{B}_{ext} = B_{ext} \hat{z}$:

$$H_D (|l, m_l\rangle |1/2, m_s\rangle) = \frac{e_0 B_{ext}}{2m} (\hat{L}_z + 2\hat{S}_z) |l, m_l\rangle |1/2, m_s\rangle$$

$$= \frac{e_0 \hbar}{2m} B_{ext} (m_l + 2m_s) |l, m_l\rangle |1/2, m_s\rangle$$

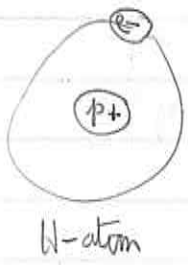
μ_B "Bohr magneton"



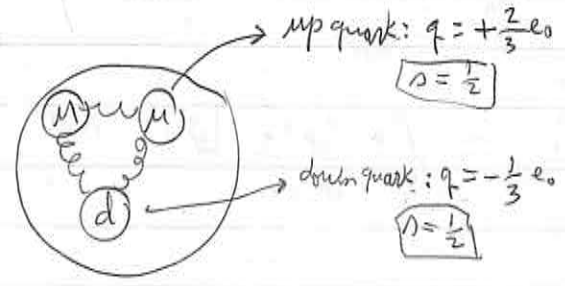
Intermediate field case: Need to diagonalize!

$$B_{ext} \approx B_{int}$$

Hyperfine effect



Nuclear spin of protons?



Add three spins - 1/2: $\frac{1}{2} \otimes \frac{1}{2} \otimes \frac{1}{2} = ?$

$\frac{1}{2} \otimes \frac{1}{2} = 1 \oplus 0$

$\frac{1}{2} \otimes \frac{1}{2} \otimes \frac{1}{2} = (1 \oplus 0) \otimes \frac{1}{2} = (1 \otimes \frac{1}{2}) \oplus (0 \otimes \frac{1}{2})$

$= \left(\frac{3}{2} \oplus \frac{1}{2} \right) \oplus \frac{1}{2}$
 $\downarrow \quad \downarrow \quad \downarrow$
 $2 \times \frac{3}{2} + 1 = 4 \quad \frac{1}{2} \quad \frac{1}{2} \quad \dim = 8 \checkmark$

⇒ Proton can have nuclear spin $\frac{3}{2}$ or $\frac{1}{2}$ ⇒ $\frac{1}{2}$ is lower energy! (Need gamma ray to excite $\frac{3}{2}$ manifold...)

ps: It's not as simple... See R. Jaffe, "When does the proton really get its spin?" Physics Today 48, 24 (1995).

Because proton has spin, it produces a magnetic field on the electron:

Proton's magnetic moment: $\vec{\mu}_p = \frac{g_p e_0}{2m_p} \vec{S}_p \quad (g_p = 5.58)$

Produces a dipole field:

$$\vec{B}_{\text{atom}} = \frac{\mu_0}{4\pi r^3} \left[3(\vec{\mu}_p \cdot \hat{n})\hat{n} - \vec{\mu}_p \right] + \frac{2}{3}\mu_0 \vec{\mu}_p \delta(\vec{r})$$

↑ $\delta(\vec{r})$ required: You can get that by calculating the field of a rotating charged sphere.

The Hamiltonian of the electron in the proton's dipole moment is

$$H_{\text{hyp}} = -\vec{\mu}_e \cdot \vec{B} = -\left(\frac{-e_0}{m_e} \vec{S}_e\right) \cdot \left[\frac{\mu_0}{4\pi r^3} (3\vec{\mu}_p \cdot \hat{n})\hat{n} - \vec{\mu}_p \right] + \frac{2\mu_0}{3} \vec{\mu}_p \delta(\vec{r})$$

$$H_{\text{hyp}} = \frac{\mu_0 g_p e_0^2}{8\pi m_p m_e} \left[\frac{3(\vec{S}_p \cdot \hat{n})(\vec{S}_e \cdot \hat{n}) - \vec{S}_p \cdot \vec{S}_e}{r^3} \right] + \frac{\mu_0 g_p e_0^2}{3 m_p m_e} (\vec{S}_p \cdot \vec{S}_e) \delta(\vec{r})$$

Calculate 1st order correction:

$$E_{\text{hyperfine}}^1 = \langle H_{\text{hyp}} \rangle = \frac{\mu_0 g_p e_0^2}{8\pi m_p m_e} \left\langle \frac{3(\vec{S}_p \cdot \hat{n})(\vec{S}_e \cdot \hat{n}) - \vec{S}_p \cdot \vec{S}_e}{r^3} \right\rangle + \frac{\mu_0 g_p e_0^2}{3 m_p m_e} \langle \vec{S}_p \cdot \vec{S}_e \rangle |\psi(0)|^2$$

$$\int \psi^* H_{\text{hyp}} \psi d^3r$$

In the ground state of the H-atom, the wave function is spherically symmetric:

choose $\hat{n} = \hat{z}$

$$\left\langle \frac{3(\vec{S}_p \cdot \hat{n})(\vec{S}_e \cdot \hat{n}) - \vec{S}_p \cdot \vec{S}_e}{r^3} \right\rangle = \left\langle \frac{3 S_{pz} S_{ez} (\cos^2 \theta) + S_{ex} S_{px} \sin^2 \theta + S_{ey} S_{py} \sin^2 \theta + S_{ez} S_{ez} \cos^2 \theta}{r^3} \right\rangle - \left\langle \frac{1}{r^3} \right\rangle S_{pz} S_{ez}$$

only the first term survives isotropic average!

$$= \left\langle \frac{3}{r^3} \right\rangle S_{pz} S_{ez} \frac{1}{4\pi} \int_0^{2\pi} d\phi \int_0^\pi \sin \theta \cos^2 \theta d\theta - \left\langle \frac{1}{r^3} \right\rangle S_{pz} S_{ez}$$

$$= \left\langle \frac{3}{r^3} \right\rangle S_{pz} S_{ez} \frac{1}{4\pi} \frac{4\pi}{3} - \left\langle \frac{1}{r^3} \right\rangle S_{pz} S_{ez} = 0!$$

Since for spherically symmetric ground state:

$$E_{\text{hyperfine}}^1 = \frac{\mu_0 g_p e_0^2}{3 m_p m_e} \langle \vec{S}_p \cdot \vec{S}_e \rangle |\psi_{100}(0)|^2 = \frac{\mu_0 g_p e_0^2}{3\pi m_p m_e a_B^3} \langle \vec{S}_p \cdot \vec{S}_e \rangle$$

"Spin-spin" coupling, instead of "spin-orbit".

How to diagonalize? Use

$$\vec{S} = \vec{S}_e + \vec{S}_p \Rightarrow \vec{S}_e \cdot \vec{S}_p = \frac{1}{2} (S^2 - S_e^2 - S_p^2) = \frac{1}{2} (S^2 - 2 \times \frac{3}{4} \hbar^2)$$

$\begin{matrix} \underbrace{\hspace{2cm}}_{= \frac{3}{4} \hbar^2} & \underbrace{\hspace{2cm}}_{= \frac{3}{4} \hbar^2} & \downarrow \\ & & = 1(1+1)\hbar^2 \text{ for Triplet} \\ & & = 0 \text{ for singlet} \end{matrix}$

$$\frac{1}{2} \otimes \frac{1}{2} = 1 \oplus 0$$

$\begin{matrix} \downarrow & \downarrow \\ \text{Triplet} & \text{Singlet} \end{matrix}$

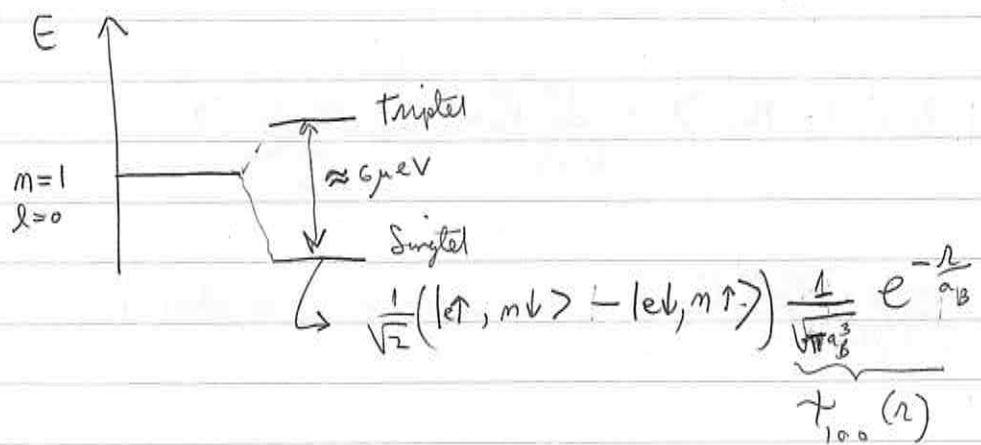
$$= \begin{cases} +\frac{1}{4} \hbar^2 \text{ for Triplet} \\ -\frac{3}{4} \hbar^2 \text{ for Singlet} \end{cases}$$

$$\Rightarrow E_{\text{hyf}}^1 = \frac{\mu_0 g_p e^2 \hbar^2}{3\pi m_p m_e a_B^3} \begin{cases} +\frac{1}{4}, \text{ Triplet} \\ -\frac{3}{4}, \text{ Singlet} \end{cases} = \frac{4 g_p \hbar^4}{3\pi m_p m_e^2 c^2 a_B^4} \begin{cases} +\frac{1}{4} \text{ (Triplet)} \\ -\frac{3}{4} \text{ (Singlet)} \end{cases}$$

$$= \frac{1}{c^2} \frac{4}{\pi} \frac{m_e e^2 \hbar^2}{4 \hbar^2} \frac{\hbar^2}{m_e e^2} \frac{\mu_0 g_p \hbar^2}{3\pi m_p m_e a_B^3}$$

$$= \frac{4}{3c^2} \frac{\hbar^4}{m_e m_p} \frac{g_p}{a_B^4} = 5.88 \times 10^{-6} \text{ eV}$$

Corresponding wavelength: $E_{\text{hyf}}^1 = h\nu = \frac{hc}{\lambda} \Rightarrow \lambda = \frac{hc}{E_{\text{hyf}}^1} = 21 \text{ cm} //$



Most of the universe is made up of Hydrogen, and it's cold ($T \approx 3 \text{ K}$) \Rightarrow The 21 cm line shows up everywhere we look!